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Explicit bounds for electromagnetic transmission problems

Andrea Moiola



Joint work with E.A. Spence (Bath)

Maxwell equations in heterogeneous media

Given:

- wavenumber k > 0
- \blacktriangleright sources $\mathbf{J},\mathbf{K}\in H(\operatorname{div}^0;\mathbb{R}^3)$, compactly supported
- $\blacktriangleright \ \epsilon_0, \mu_0 > 0$
- $ightarrow \epsilon, \mu \in L^\infty(\mathbb{R}^3;SPD)$ such that

 $\Omega_{\mathbf{i}} := \operatorname{int}(\operatorname{supp}(\epsilon - \epsilon_0 \underline{\underline{\mathbf{I}}}) \cup \operatorname{supp}(\mu - \mu_0 \underline{\underline{\mathbf{I}}})) \text{ is bounded and Lipschitz}$

Find $\mathbf{E}, \mathbf{H} \in H_{\text{loc}}(\text{curl}; \mathbb{R}^3)$ such that $\mathbf{i} k \epsilon \mathbf{E} + \nabla \times \mathbf{H} = \mathbf{J}$ in \mathbb{R}^3 , $-\mathbf{i} k \mu \mathbf{H} + \nabla \times \mathbf{E} = \mathbf{K}$ in \mathbb{R}^3 , (\mathbf{E}, \mathbf{H}) satisfy Silver–Müller radiation condition $|\sqrt{\epsilon_0} \mathbf{E} - \sqrt{\mu_0} \mathbf{H} \times \frac{\mathbf{x}}{|\mathbf{x}|}| = \mathcal{O}_{|\mathbf{x}| \to \infty}(|\mathbf{x}|^{-2}).$ $\epsilon = \epsilon_0$ $\mu = \mu_0$

Special case: "transmission problem", i.e. homogeneous scatterer

$$\epsilon = \begin{cases} \epsilon_i \\ \epsilon_0 \end{cases} \quad \mu = \begin{cases} \mu_i & \text{ in } \Omega_i \\ \mu_0 & \text{ in } \Omega_o := \mathbb{R}^3 \setminus \overline{\Omega_i} \end{cases} \quad 0 < \epsilon_i, \epsilon_0, \mu_i, \mu_0 \text{ constant.} \end{cases}$$

Wave scattering

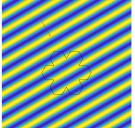
The example we have in mind is incident wave \mathbf{E}^{Inc} , \mathbf{H}^{Inc} hitting Ω_i :

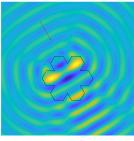
 \rightarrow BVP with data supported on Ω_i :

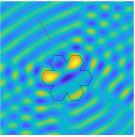
Incoming field

$$\mathbf{J}=\mathrm{i}k^2(\epsilon_0-\epsilon)\mathbf{E}^{Inc}$$
, $\mathbf{K}=\mathrm{i}k^2(\mu-\mu_0)\mathbf{H}^{Inc}$,

Total field Scattered field
$$\begin{split} \mathbf{E}^{Inc} &= \sqrt{\frac{\mu_0}{\epsilon_0}} \mathbf{A} e^{ik\sqrt{\epsilon_0\mu_0}\mathbf{x}\cdot\mathbf{d}} \\ \mathbf{H}^{Inc} &= \mathbf{d} \times \mathbf{A} e^{ik\sqrt{\epsilon_0\mu_0}\mathbf{x}\cdot\mathbf{d}} \end{split}$$
 $\mathbf{E} + \mathbf{E}^{Inc}$ Е $\mathbf{H} + \mathbf{H}^{Inc}$ н







datum

BVP solution

physical field

If ϵ,μ are sufficiently regular then the problem is well-posed. From Fredholm theory we have

$$\left\| \left(\begin{array}{c} \mathbf{E} \\ \mathbf{H} \end{array}\right) \right\|_{\Omega_{i/o}} \leq \mathbf{C} \left\| \left(\begin{array}{c} \mathbf{J} \\ \mathbf{K} \end{array}\right) \right\|_{\Omega_{i/o}}$$

Goal: find out how $C = C(k, \epsilon, \mu)$ depends on k, ϵ and μ .

Why? In FEM & BEM analysis and in UQ for time-harmonic problems, explicit parameter dependence allows to control:

- Quasi-optimality & pollution effect
- Gmres iteration numbers
- Matrix compression
- hp-FEM & BEM (Melenk–Sauter)
- Shape differentiation & uncertainty quantification



LAFONTAINE, SPENCE, WUNSCH, arXiv 2019:

(Helmholtz)

The following is a non-exhaustive list of papers on the frequency-explicit convergence analysis of numerical methods for solving the Helmholtz equation where a central role is played by *either* the non-trapping resolvent estimate (1.5), *or* its analogue (with the same *k*-dependence) for the commonly-used approximation of the exterior problem where the exterior domain \mathcal{O}_+ is truncated and an impedance boundary condition is imposed:

- conforming FEMs (including continuous interior-penalty methods) [72, Proposition 2.1], [74, Proposition 8.1.4], [56, Lemma 2.1], [77, Lemma 3.5], [78, Assumptions 4.8 and 4.18], [45, §2.1], [110, Theorem 3.1], [113, §3.1], [44, §3.2.1], [40, Remark 3.2], [41, Remark 3.1], [29, Assumption 1], [30, Definition 2], [55, Theorem 3.2], [50, Lemma 6.7], [14, Equation 4],
- least squares methods [33, Assumption A1], [10, Remark 1.2], [64, Assumption 1 and equation after Equation 5.37],
- DG methods based on piece-wise polynomials [46, Theorem 2.2], [47, Theorem 2.1], [39, Assumption 3], [48, §3], [62, Assumption A (Equation 4.5)], [76, Equation 4.4], [35, Remark 3.2], [32, Equation 2.4], [84, Equation 4.3], [112, Remark 3.1], [94, Theorem 2.2],
- plane-wave/Trefftz-DG methods [3, Theorem 1], [59, Equation 3.5], [60, Theorem 2.2], [2, Lemma 4.1], [61, Proposition 2.1],
- multiscale finite-element methods [51, Equation 2.3], [13, §1.2], [88, Assumption 5.3], [8, Theorem 1], [87, Assumption 3.8], [31, Assumption 1],
- integral-equation methods [71, Equation 3.24], [75, Equation 4.4], [24, Chapter 5], [53, Theorem 3.2], [111, Remark 7.5], [43, Theorem 2], [49, Theorem 3.2], [52, Assumption 3.2],

In addition, the following papers focus on proving bounds on the solution of Helmholtz boundary-value problems (with these bounds often called "stability estimates") motivated by applications in numerical analysis: [36], [57], [26], [11], [7], [70], [98], [28], [6], [9], [27], [93], [54], [55] [83], [50], Of these papers, all but [70], [6], [27], [11] are in nontrapping situations, [70], [6], [27] are in parabolic trapping scenarios, and [11] proves the exponential growth (1.7) under elliptic trapping.

What about Helmholtz?

Simplest heterogeneous Helmholtz problem: find $u \in H^1_{\mathrm{loc}}(\mathbb{R}^d)$ s.t.

 $\begin{array}{ll} \Delta u + k^2 \, n \, u = f & \text{ in } \mathbb{R}^d \\ \text{+Sommerfeld radiation c.} & f \in L^2(\mathbb{R}^d), \quad n = \begin{cases} n_i & \text{ constant in } \Omega_i, \\ 1 & \text{ in } \Omega_o. \end{cases}$

▶ If $0 < n_i < 1$, Ω_i star-shaped

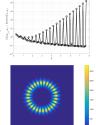
 $\Omega_i \cup \operatorname{supp} f \subset B_R$

$$\left\|
abla u
ight\|_{L^{2}(B_{R})}^{2} + k^{2} \left\| \sqrt{n} \, u
ight\|_{L^{2}(B_{R})}^{2} \leq \left[4R^{2} + rac{1}{n_{i}} \left(2R + rac{d-1}{k}
ight)^{2}
ight] \left\| f
ight\|_{L^{2}(B_{R})}^{2}$$

Fully explicit, k-independent, shape-robust estimate. (For d = 2 it implies bounds for Maxwell TE/TM modes.)

▶ If $n_i > 1$, Ω_i strictly convex & C^∞ : superalgebraic blow up in k, quasi-resonances, ray trapping, creeping waves...

Dependence on parameters is complicated! Monotonicity of n & shape of Ω_i are crucial.



Wavenumber-explicit bounds: a bit of history

MORAWETZ 1960s/70)s: introd	uced main tools (multipliers)
Melenk 1995:	1st <i>k</i> -explicit bound for Helmholtz, bdd dom.	
CHANDLER-WILDE, MONK 2008:		unbounded domains
► HIPTMAIR, MOIOLA, PERUGIA 2011:		Maxwell, bdd dom.
homogeneous coeff. heterogeneous coeff.		
heterogeneous coeff.		
► MOIOLA, SPENCE 201	9: Helmh	oltz & piecewise-constant n
► Graham, Pembery, Spence 2019:		Helmholtz & general coeff.
VERFÜRTH 2019:		Maxwell & impedance

Plenty of other related contributions exist! BARUCQ, CHAUMONT-FRELET, FENG, HETMANIUK, LORTON, PETERSEIM, SAUTER, TORRES, WIENERS&WOHLMUTH, (your name here), ...

Our goal: extend (GRAHAM, PEMBERY, SPENCE 2019) to Maxwell eq.s.

Bound #1: transmission problem

Single homogeneous scatterer:

$$\begin{split} \epsilon &= \begin{cases} \epsilon_i & \text{in } \Omega_i \\ \epsilon_0 & \text{in } \Omega_o \end{cases}, \quad \mu = \begin{cases} \mu_i & \text{in } \Omega_i \\ \mu_0 & \text{in } \Omega_o \end{cases} \quad 0 < \epsilon_i, \epsilon_0, \mu_i, \mu_0 \text{ constant.} \end{cases} \\ \mathbf{f} \underbrace{\epsilon_i \leq \epsilon_0}_{i_1}, \underbrace{\mu_i \leq \mu_0}_{i_1}, \underbrace{\Omega_i \text{ star-shaped}}_{i_1}, \Omega_i \cup \text{supp } \mathbf{J} \cup \text{supp } \mathbf{K} \subset B_R, \text{ then} \end{cases} \\ \underbrace{\epsilon_i \|\mathbf{E}\|_{B_R}^2 + \mu_i \|\mathbf{H}\|_{B_R}^2}_{i_1 \in \mathbb{C}} &\leq 4R^2 \Big(\frac{\epsilon_0}{\epsilon_i} + \frac{\mu_0}{\mu_i}\Big) \Big(\epsilon_0 \|\mathbf{K}\|_{B_R}^2 + \mu_0 \|\mathbf{J}\|_{B_R}^2 \Big). \\ & \|\cdot\|_{B_R} = \|\cdot\|_{L^2(B_R)} \end{cases} \\ \end{bmatrix} \\ \text{Equivalent to wavenumber-independent } H(\text{curl}; B_R) \text{ bound for } \mathbf{E}. \end{split}$$

If ϵ_i is (constant) SPD matrix, same holds if $\max eig(\epsilon_i) \le \epsilon_0$ and with ϵ_i substituted by $\min eig(\epsilon_i)$ in the bound. Same for μ_i .

Bound #2: more general ϵ, μ

Assume $\epsilon, \mu \in W^{1,\infty}(\Omega_i; SPD)$, Ω_i Lipschitz,

 $\blacktriangleright \Omega_i$ star-shaped

 $\blacktriangleright \|\epsilon_i\|_{L^{\infty}(\partial\Omega_i)} \leq \epsilon_0, \ \|\mu_i\|_{L^{\infty}(\partial\Omega_i)} \leq \mu_0, \ \text{ i.e. jumps are "upwards" on } \partial\Omega_i$

- ► $\epsilon_* := \operatorname{ess\,inf}_{\mathbf{x}\in\Omega_l} \left(\epsilon + (\mathbf{x}\cdot\nabla)\epsilon\right) > 0, \ \mu_* := \operatorname{ess\,inf}_{\mathbf{x}\in\Omega_l} \left(\mu + (\mathbf{x}\cdot\nabla)\mu\right) > 0$ "weak monotonicity" in radial direction, avoid trapping of rays
- "extra regularity" ($\mathbf{E}, \mathbf{H} \in H^1(\Omega_i \cup \Omega_o)^3$ or $\epsilon, \mu \in C^1(\Omega_i)$ or $W^{1,\infty}(\mathbb{R}^3)$)

Then we have explicit wavenumber-independent bound:

$$\begin{split} \epsilon_* \left\| \mathbf{E} \right\|_{B_R}^2 + \mu_* \left\| \mathbf{H} \right\|_{B_R}^2 \\ &\leq 4R^2 \bigg(\frac{\left\| \epsilon \right\|_{L^{\infty}(B_R)}^2}{\epsilon_*} + \frac{\epsilon_0 \mu_0}{\mu_*} \bigg) \left\| \mathbf{K} \right\|_{B_R}^2 + 4R^2 \bigg(\frac{\left\| \mu \right\|_{L^{\infty}(B_R)}^2}{\mu_*} + \frac{\epsilon_0 \mu_0}{\epsilon_*} \bigg) \left\| \mathbf{J} \right\|_{B_R}^2. \end{split}$$

Expect (from Helmholtz analogy) superalgebraic blow up in k if any of the first 3 assumptions is lifted.

Similar results when \mathbb{R}^3 is truncated with impedance BCs.

How our bound was obtained

First consider smooth case $\mathbf{E}, \mathbf{H} \in \mathbf{C}^1(\mathbb{R}^3; \mathbb{C}^3)$.

(i) Multiply the 2 PDEs by the "test fields" (Morawetz multipliers)

$$\begin{array}{ll} (\epsilon \overline{\mathbf{E}} \times \mathbf{x} + R\sqrt{\epsilon \mu} \overline{\mathbf{H}}) & \& & (\mu \overline{\mathbf{H}} \times \mathbf{x} - R\sqrt{\epsilon \mu} \overline{\mathbf{E}}) & \quad \text{in } B_R \supset \Omega_i, \\ \epsilon_0 \overline{\mathbf{E}} \times \mathbf{x} + r\sqrt{\epsilon_0 \mu_0} \overline{\mathbf{H}}) & \& & (\mu_0 \overline{\mathbf{H}} \times \mathbf{x} - r\sqrt{\epsilon_0 \mu_0} \overline{\mathbf{E}}) & \quad \text{in } \mathbb{R}^3 \setminus B_R, \end{array}$$

(ii) integrate by parts in Ω_i , $B_R \setminus \overline{\Omega_i}$ and $\mathbb{R}^3 \setminus B_R$, (iii) sum 3 contributions, (iv) take \Re eal part, (v) have fun!

$$\begin{split} &\int_{B_R} \overline{\mathbf{E}} \cdot \underbrace{\left(\boldsymbol{\epsilon} + (\mathbf{x} \cdot \nabla) \boldsymbol{\epsilon} \right)}_{\geq \boldsymbol{\epsilon}_* \text{ by assumpt.}} \mathbf{E} + \overline{\mathbf{H}} \cdot \underbrace{\left(\boldsymbol{\mu} + (\mathbf{x} \cdot \nabla) \boldsymbol{\mu} \right)}_{\geq \boldsymbol{\mu}_* \text{ by assumpt.}} \mathbf{H} & \underset{\nabla \cdot [\boldsymbol{\epsilon} \mathbf{E}] = \nabla \cdot [\boldsymbol{\mu} \mathbf{H}] = 0}{\overset{\text{Using PDEs } \&}{\nabla \cdot [\boldsymbol{\epsilon} \mathbf{E}] = \nabla \cdot [\boldsymbol{\mu} \mathbf{H}] = 0}} \\ &= 2 \int_{B_R} \Re \Big\{ \mathbf{K} \cdot \left(\boldsymbol{\epsilon} \overline{\mathbf{E}} \times \mathbf{x} + \sqrt{\boldsymbol{\epsilon}_0 \boldsymbol{\mu}_0} R \overline{\mathbf{H}} \right) + \mathbf{J} \cdot \left(\boldsymbol{\mu} \overline{\mathbf{H}} \times \mathbf{x} - \sqrt{\boldsymbol{\epsilon}_0 \boldsymbol{\mu}_0} R \overline{\mathbf{E}} \right) \Big\} \\ &+ \int_{\partial \Omega_l} \underbrace{\left[\text{terms from IBP} \right]}_{[\mathbf{E}_T, \mathbf{H}_T, (\boldsymbol{\epsilon} \mathbf{E})_N, (\boldsymbol{\mu} \mathbf{H})_N] = 0} + \int_{\partial B_R} \underbrace{\left[\text{terms from IBP} \right]}_{\leq \mathbf{0} \text{ by S-M radiation c.}} \end{split}$$

Conclude by Cauchy-Schwarz.

Rough coefficients, regularity and density

Proof in previous slide only uses elementary results if ${f E}, {f H} \in C^1({\mathbb R}^3; {\mathbb C}^3).$

For general case we need density of inclusion

$$C^{\infty}(\overline{D})^{3} \subset \left\{ \mathbf{v} \in H(\operatorname{curl}; D), \nabla \cdot [\mathsf{A}\mathbf{v}] \in L^{2}(D), \mathsf{A}\mathbf{v} \cdot \hat{\mathbf{n}} \in L^{2}(\partial D), \mathbf{v}_{T} \in L^{2}_{T}(\partial D) \right\}$$

for $A = \epsilon \& A = \mu$, D Lipschitz bdd.

If $A \in C^1(\Omega_i; SPD)$, this density is non-trivial but follows from regularity results for layer potentials on manifolds (MITREA, TAYLOR 1999).

- Equivalent step for Helmholtz was much simpler.
- Constant scalar $\epsilon \& \mu$: density proved in COSTABEL, DAUGE 1998.
- ▶ If $\mathbf{E}, \mathbf{H} \in H^1_{\text{loc}}(\mathbb{R}^3; \mathbb{C}^3)$ then no density is needed. E.g. ensured if $\epsilon, \mu \in W^{1,\infty}(\mathbb{R}^3; SPD)$ (no jumps).
- What about $A \in W^{1,\infty}(\Omega_i; SPD)$?

Summary

Time-harmonic Maxwell eq.s in \mathbb{R}^3 with heterogeneous inclusion:

- ► fully explicit bounds on $\|\mathbf{E}\|_{H(\operatorname{curl},B_R)}$ if ϵ, μ "radially growing"
- ▶ also for impedance BVPs in star-shaped domains
- ▶ extends Helmholtz results from (GRAHAM, PEMBERY, SPENCE 2019)

Some open questions:

- resonance-free strip in complex k plane?
- ▶ presence of quasi-resonances blow up for "wrong" coefficients?
- ▶ rougher ($W^{1,\infty}(\Omega_i; SPD), L^{\infty}$) coefficients?
- ▶ relation with shape-differentiation and UQ?

Preprint coming soon...

